

# A GENERALISED $(M, N_R)$ MIMO RAYLEIGH CHANNEL MODEL FOR NON-ISOTROPIC SCATTERER DISTRIBUTIONS

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## ABSTRACT

This paper extends a recently proposed space-time model for Rayleigh fading to include an arbitrary transmit antenna configuration of any shape and size transmitting simultaneously in a multiple-input multiple-output (MIMO) channel. The space-time correlation function and space-frequency cross spectrum function for a non-isotropic scatterer distribution around the receiver is derived for the arbitrary configuration as a further extension of a previous result of a multiple-input single-output Rayleigh wireless channel which used a ring of uniformly distributed scatterers model. Analysis based on achievable spectral efficiency for typical arbitrary transmit antenna configurations is given. The analysis demonstrates the utility of the correlation function.

## INTRODUCTION

The combination of temporal and spatial diversity are two effective means by which communication quality and associated system performance can be significantly improved in a rich-scattering wireless environment. This has been demonstrated through proposals for space-time coded modulation over multiple-input multiple-output (MIMO) radio channels which are understood to be beneficial for high data-rate systems operating in rich-scattering wireless environments. Without designing highly detailed system models, which can be used for individual channel realisations, it is possible to obtain a more rigorous evaluation of the proposed coding schemes through simpler macroscopic system models. One such scheme, based on a traditional Clarke/Jakes model [1], has been proposed in [2], where a rich isotropic distribution of scatterers is assumed around the mobile station (MS), the receiver. No major scatterers are located around the base station (BS), the transmitter. Space-time correlation and space-frequency cross spectrum functions corresponding to this distribution were derived. This was extended to the case of a  $(2, N_R)$ , i.e. 2 transmit,  $N_R$  receive antennas, Rayleigh fading channel with a non isotropic distribution of scatterers in [3].

One major drawback of [2, 3] is that by consideration of time selectivity and/or frequency selectivity (with reasonable sized Doppler spread, and/or delay spread), limits the models to  $M=2$  transmit antennas at a particular time instant. Thus, they are not applicable to many proposed mobile communication systems where application of space-time coding utilises  $M > 2$  for continuous fast fading and/or frequency selective fading wireless scenarios. This paper will provide a useful generalisation of [3] that will mean the space-time model can be applied to any numbers of antennas at the BS in a macrocellular fast fading scenario. The application of the model can be for frequency selective and/or frequency non-selective wireless scenarios through derivation of appropriate space-time cross correlation, and space-frequency cross spectrum functions. Such functions can analyse space-time coding, or space-frequency coding, schemes. To demonstrate the utility of the proposed model a simple spectral efficiency analysis based on Shannon's capacity measure is given for the case of a  $(4, 4)$ , fast flat Rayleigh fading wireless channel.

## SPACE-TIME CROSS CORRELATION FORMULATION

Consider an arbitrary number of BS antennas,  $M$ , of any configuration, located at  $y_1, y_2, \dots, y_M$ , with no major scatterers and only characterised by horizontal separation of an arbitrary number of wavelengths ( $\lambda$ ), and  $N_R$  approximately co-located receive antennas at the MS. The discussion in this paper assumes a macrocellular radio channel. Fig. 1 illustrates the MIMO transmission model between the BS and MS for the case of  $M = 4$ .

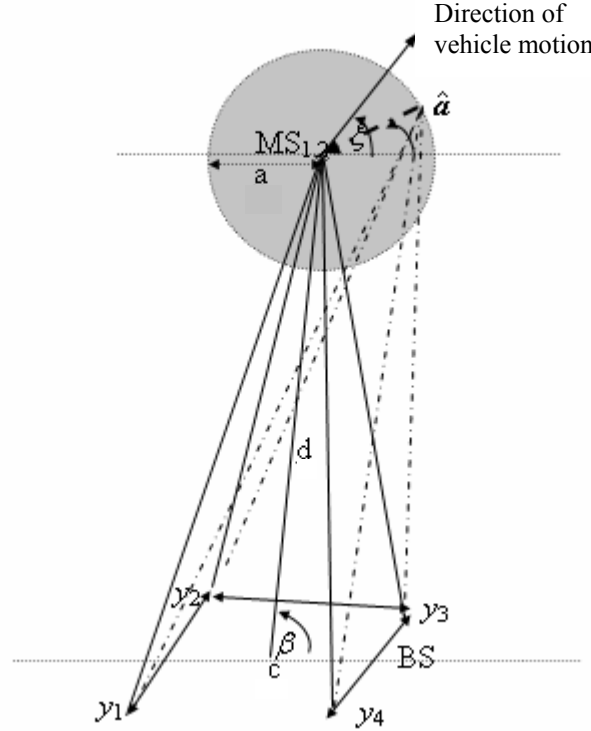


Fig. 1 (4, 2),  $M = 4$ ,  $N_R = 2$ , MIMO Transmission model between base station (BS) and mobile station (MS).

The space-time cross correlation function can be defined as  $\rho(y_{m''}, y_{m'}; \tau) \triangleq E c_{m'}(t) c_{m''}^*(d_{sp}, t - \tau) = \rho(d_{sp}, \tau)$ , and in a manner similar to [2] we obtain,

$$\rho(d_{sp}, \tau) = \sigma^2 e^{(j k d_{sp} \cos(\beta - \gamma))} \times \int_{\Omega} P(\hat{a}) \left( e^{[j \cos \hat{a} (2\pi f_D \tau \cos \xi + k z_{cs}')] } \times e^{[j \sin \hat{a} (2\pi f_D \tau \sin \xi - k z_{ss}')] } \right) d\hat{a} \quad (1)$$

where  $d_{sp} = |y_{m''} - y_{m'}|$  is the magnitude separation of antennas  $m''$  and  $m'$  at the BS,  $k = 2\pi/\lambda$  is the wave number,  $\hat{a}$  is a unit vector pointing in the direction of wave propagation with reference to the MS,  $\Omega$  corresponds to the unit circle over which integration is performed,  $P(\hat{a})$  is the angular power density distribution function around the MS of scatterers and  $j = \sqrt{-1}$ , and  $\gamma$  is the angle of separation between antennas  $y_{m''}$  and  $y_{m'}$  with respect to  $y_1$ . Note that for uniformly distributed scatterers  $P(\hat{a}) = 1$  and for a finite number of scatterers the integration in (1) reduces to a summation as in [2]. Also with reference to (1),  $f_D$  is the maximal Doppler spread,  $\xi$  is the direction of movement of the MS with reference to the BS,  $\sigma$  is the variance of the channel gain, and  $z_{cs}'$  and  $z_{ss}'$ , found using macrocellular far-field assumptions [2], are similar to those defined in [2, App. 1, eqn. (36,37)] as

$$z_{cs}' = c_s \sin \beta, \quad z_{ss}' = c_s \cos \beta \quad (2)$$

where  $c_s = d_{sp} \sin(\beta - \gamma) \times a/d'$ ,  $a/d'$  represents the ratio of distance,  $a$ , of scatterers from the MS and the distance,  $d'$ , of the MS to the centre of the antennas  $m''$  and  $m'$ , and  $\beta$  represents the angular position of the MS with respect to the BS. To further appreciate the material in this section the reader is referred to [2].

Similarly to [2], and using the 2-D modal expansion as in [4], if one lets  $\hat{\mathbf{a}} = (1, \phi)$  the following formulation is obtained for the space-time correlation function

$$\rho(d_{sp}, \tau) \equiv R_{c_m c_{m'}}(d_{sp}, \tau) = \sigma^2 e^{jkd_{sp} \cos(\beta - \gamma)} \cdot \int_0^{2\pi} \sum_{m=-\infty}^{\infty} \gamma_m e^{-jm\phi} e^{j \left( \cos \phi (2\pi f_D \tau \cos \xi + kz_{cs}') + \sin \phi (2\pi f_D \tau \sin \xi - kz_{ss}') \right)} d\phi \quad (3)$$

After an appropriate change of variables the following is obtained for the space-time correlation function,

$$R_{c_m c_{m'}}(\tau, d_{sp}) = \sigma^2 e^{jkd_{sp} \cos(\beta - \gamma)} \cdot 2\pi \sum_{m=-\infty}^{\infty} \gamma_m e^{jm\psi} J_m(z). \quad (4)$$

where  $\psi = \tan^{-1}(b_1/a_1)$  and  $z = 2\pi\sqrt{a_1^2 + b_1^2}$ ;  $b_1 = (f_D \tau \sin \xi - z_{ss}')/\lambda$  and  $a_1 = (f_D \tau \cos \xi + z_{cs}')/\lambda$ ; and  $J_m(\cdot)$  is the  $m$ th order Bessel function of the first kind.

Note that for an isotropic scatterer distribution,

$$\gamma_m = \begin{cases} 1/2\pi, & m = 0 \\ 0, & m \neq 0 \end{cases}, \text{ and } R_{c_m c_{m'}}(d_{sp}, \tau) = \sigma^2 e^{jkd_{sp} \cos(\beta - \gamma)} J_0(z) \quad (5)$$

, which for the case of  $M = 2$ , and thus  $\beta = \gamma$ , we obtain the same space-time correlation as defined in [4].

A 3-D plot of the magnitude of the cross-correlation function,  $|R_{c_m c_{m'}}(d_{sp}, \tau)|$ , is shown in Fig. 2 for  $f_D = 0.01$ , between diagonally opposite antennas for a uniform circular array (UCA),  $M = 4$ , with antenna spacing,  $d_{sp} = 10\lambda$  (4 element UCA also shown in Fig. 2);  $\xi = \pi/3$  and  $\beta = \pi/6$  assuming a Laplacian distribution [5]. The angular spread,  $S_\sigma^2$ , of  $10^\circ$  given, is found from the square root of the variance, [4].  $|R_{c_m c_{m'}}(d_{sp}, \tau)|$  can also be plotted for all other common angular power distributions for BS antenna configurations of arbitrary shape and size.

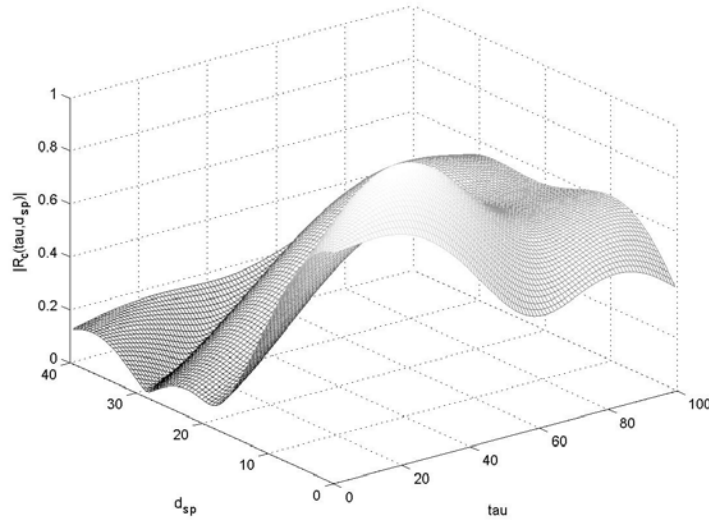


Fig. 2 3-D plot of magnitude of cross-correlation function,  $|R_c(d_{sp}, \tau)|$ , for  $f_D = 0.01$ , MS moving direction with respect to BS,  $\xi = \pi/3$ , and MS position with respect to BS,  $\beta = \pi/6$ , assuming a Laplacian distribution with an angular spread of  $10^\circ$ .

## SPACE-FREQUENCY CROSS SPECTRUM FORMULATION

Following from (5), and in a manner similar to the derivation of the space-frequency cross spectrum given in [4, eqn (3), p. 1176], a space-frequency cross spectrum can be found for a non-isotropic distribution of scatterers. We first observe that  $z$  can be reformulated as

$$z = \sqrt{(2\pi[f_D\tau + c_s \sin(\beta - \xi)])^2 + (2\pi c_s \cos(\beta - \xi))^2}. \quad (6)$$

The space-frequency cross spectrum is defined as  $S_{c_m c_m^*}(d_{sp}, f) \triangleq \mathcal{F}\{R_{c_m c_m^*}(\tau, d_{sp})\}$  where  $\mathcal{F}\{\cdot\}$  is the Fourier transform with respect to  $\tau$ . Thus

$$S_{c_m c_m^*}(d_{sp}, f) = \sigma^2 e^{(jkd_{sp} \cos(\beta - \gamma))} . 2\pi \sum_{m=-\infty}^{\infty} (\gamma_m \mathcal{F}\{e^{im\psi} J_m(z)\}). \quad (7)$$

Thus following from some manipulation of (7), and (6),  $S_{c_1 c_2}(d_{sp}, f)$ , can be expressed as

$$\begin{aligned} S_{c_1 c_2}(d_{sp}, f) &= \sigma^2 e^{(jkd_{sp} \cos(\beta - \gamma))} . 2\pi \sum_{m=-\infty}^{\infty} \left\{ \gamma_m \right. \\ &\quad \times \left. \sum_{n=-\infty}^{\infty} j^n \left( J_{m+n}(2\pi c_s \cos(\beta - \xi)) \sum_{m'=-\infty}^{\infty} [\mathcal{F}\{J_n(2\pi f_D \tau)\} J_{n-m'}(2\pi c_s \sin(\beta - \xi))] \right) \right\} \end{aligned} \quad (8)$$

where  $\mathcal{F}\{\cdot\}$  is as defined as

$$\begin{aligned} \mathcal{F}\{J_n(2\pi f_D \tau)\} &= \pm \left( \pi f_D \sqrt{1 - (f/f_D)^2} \right)^{-1} \cos(n \sin^{-1}(f/f_D)), f < f_D \\ &= \mp f_D^n f^{-n} \sin(n\pi/2) \left( \pi f_D \sqrt{1 - (f_D/f)^2} \right)^{-1} \left( 1 + \sqrt{1 - (f/f_D)^2} \right)^{-n}, f > f_D \end{aligned} \quad (9)$$

where  $f_D$  is now the maximal Doppler spread. The values  $\pm x_1 / \mp x_2$  for  $n = -\infty \dots \infty$ , and  $n \neq 0$ , depend on whether  $n > 0$  and/or  $|n|$  even, in which case the transform is  $+x_1/-x_2$ ; otherwise, if  $n < 0$  and  $n$  is odd, one has  $-x_1/+x_2$ .

It can also be readily observed that from (8), if we are summing over  $n = -\infty \dots \infty$ , then the term  $\mp x_2$  in equation (9) can be disregarded since  $x_2$  is only non-zero for odd  $n$ ,  $n = -\infty \dots \infty$ , due to the  $\sin(n\pi/2)$  term in  $x_2$  and  $j^n \sin(n\pi/2) + j^{-n} \sin(-n\pi/2) = 0$ . Consequently (8) gives a closed form expression for the space-frequency cross spectrum with a non-isotropic distribution of scatterers.

## SPECTRAL EFFICIENCY ANALYSIS FOR A FREQUENCY FLAT FADING (4, 4) MIMO CHANNEL

This section investigates the spectral efficiency of a (4, 4), i.e.  $M = 4$ ,  $N_R = 4$  co-located receive antennas, MIMO wireless channel, for Rayleigh frequency flat fading. Two separate non-isotropic scatterer distributions are considered, a  $\cos^{2p}\phi$  distribution and a Laplacian distribution. The general channel transfer matrix,  $\mathbf{H}_{M, N_R}(\tau, d_{sp})$  is found as, similarly to [11],

$$\mathbf{H}_{M, N_R}(\tau, d_{sp}) = \frac{\hat{R}_M(\tau, d_{sp})^{1/2}}{\sigma} A_{M, N_R}, \quad \text{where} \quad (10)$$

$$\hat{R}_M(\tau, d_{sp}) = \begin{bmatrix} R_c(\tau) & R_{c_1 c_2}(\tau, d_{sp}) & \cdots & R_{c_1 c_M}(\tau, d_{sp}) \\ R_{c_2 c_1}(\tau, d_{sp}) & \ddots & \ddots & \vdots \\ \vdots & \ddots & \ddots & R_{c_{M-1} c_M}(\tau, d_{sp}) \\ R_{c_M c_1}(\tau, d_{sp}) & \cdots & R_{c_M c_{M-1}}(\tau, d_{sp}) & R_c(\tau) \end{bmatrix}. \quad (11)$$

The term  $\hat{R}(\bullet)^{1/2}$  represents the space-time correlation matrix square root,  $A_{M, N_R}$  is an  $M \times N_R$  matrix of zero mean, unit variance, complex normal i.i.d. variables. The term  $R_c(\tau)$  represent the autocorrelation function for the particular non-isotropic distribution of scatterers.

Assuming  $N_R = M$ , for a rich-scattering wireless environment, we obtain the following for the Shannon capacity,  $C$ , of the  $M$  data subchannels [13] (if it is assumed that equal power is assigned to each subchannel),

$$C = \sum_{k=1}^M \log_2 \left( 1 + \lambda_k \frac{P_{BS}/M}{\sigma_n^2} \right) \quad (12)$$

where  $P_{BS}$  is the power assigned to the  $k$ th subchannel,  $\sigma_n^2$  is the noise power and  $\lambda_k$  is the  $k$ th eigenvalue of the  $M$  data subchannels, found from the eigenvalues of  $H(\tau)H(\tau)^\dagger$  where  $(\bullet)^\dagger$  represents the matrix adjoint, or complex conjugate transpose.

Fig. 3 and Fig. 4 shows the cumulative distribution function of achievable spectral efficiency calculated according to (13) for a (4, 4) channel, found via Monte Carlo simulations after  $\tau = M = 4$  time samples for a maximal Doppler spread,  $f_D = 0.01$  for non-isotropic scatterer distributions. For the Laplacian distribution, [5], an angular spread of  $10^\circ$  is assumed and for the  $\cos^{2p}\phi$  distribution, [6], an equivalent half-power beamwidth (HPBW) of the BS antennas of  $131^\circ$  is assumed. A Uniform Circular Array (UCA) is modelled at the BS where  $\beta$  and  $\xi$  are taken from uniform distributions on  $(-\pi, \pi]$ , for a equivalent antenna spacing  $d_{sp}$ , of  $1\lambda$ ,  $5\lambda$ ,  $10\lambda$  and  $20\lambda$  for an SNR,  $P_{BS}/\sigma_n^2 = 20$  dB. Fig. 3 shows the case of  $N_R = 2$  as a dashed lines to compare with the results for the Laplacian distribution, and Fig. 4 shows the comparison of Laplacian distribution and  $\cos^{2p}\phi$  distributions.

Fig. 3 shows expected performance improvement for  $N_R = 4$  over  $N_R = 2$  receive antennas. There is also demonstrable improvement in spectral efficiency with increasing BS aperture. Fig. 4 shows that there is significant improvement in spectral efficiency for the particular Laplacian scatterer distribution over the  $\cos^{2p}\phi$  distribution for smaller BS apertures. But for the larger BS apertures, i.e. for a UCA with antenna spacing,  $d_{sp}$  of  $10\lambda$  or  $20\lambda$ , the  $\cos^{2p}\phi$  distribution shows improvement in achievable spectral efficiency over the Laplacian distribution considered.

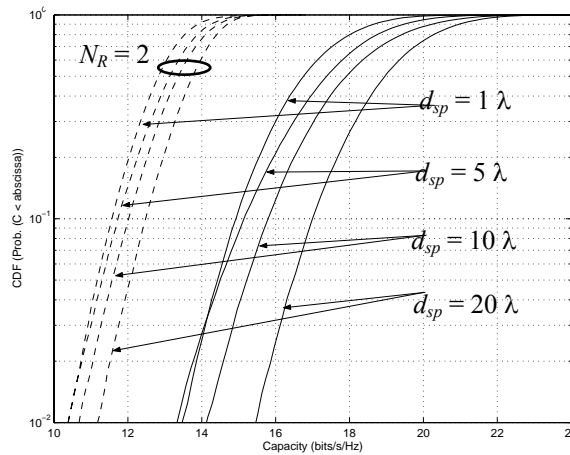


Fig. 3. Plot of cumulative distribution function (CDF) v. Capacity (bits/s/Hz) for a (4, 4) channel after  $\tau = 4$  time samples,  $f_D = 0.01$ , found via Monte Carlo simulations for a UCA at the BS with various antenna spacing,  $d_{sp}$ . A Laplacian distribution, [5], is assumed with an angular spread of  $10^\circ$ ,  $\beta$  and  $\xi$  are taken from uniform distributions on  $(-\pi, \pi]$ . The results for  $N_R = 2$  are shown as dashed lines.

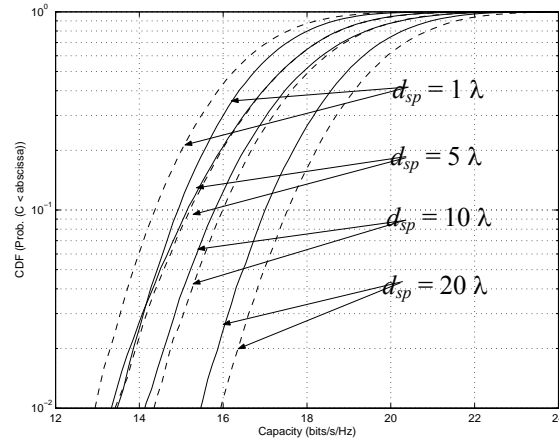


Fig. 4. Plots of CDF v. Capacity (bits/s/Hz) for a (4, 4) channel after  $\tau = 4$  time samples,  $f_D = 0.01$ , found via Monte Carlo simulations for a UCA with various antenna spacing,  $d_{sp}$ . The results for a Laplacian distribution, [5], with an angular spread of  $10^\circ$  is shown as solid lines, and those for a  $\cos^2 \phi$  distribution, [5], equivalent HPBW =  $131^\circ$ , are shown as dashed lines;  $\beta$  and  $\xi$  are taken from uniform distributions on  $(-\pi, \pi]$ .

## CONCLUDING REMARKS

Closed-form solutions have been found for the space-time cross correlation and space-frequency cross spectrum which can be applied to continuous fading  $(M, N_R)$  MIMO channels for both frequency flat and frequency selective fading wireless scenarios with non-isotropic scatterer distributions. The correlation function has been successfully applied to a Rayleigh continuous flat fading channel to obtain a measure of performance based on achievable spectral efficiency both a Laplacian distribution, and a  $\cos^2 \phi$  non-isotropic scatterer distribution.

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